

Ques I

(a) (i)

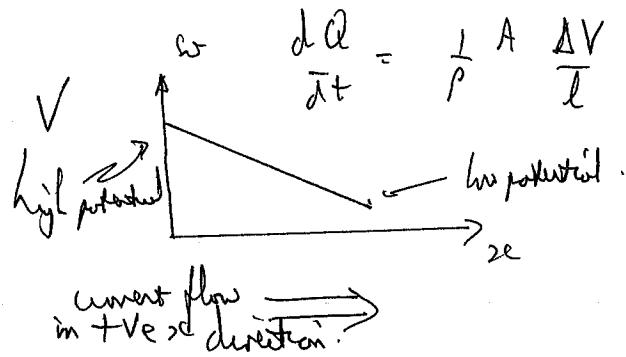
Heat
Charge

rate
 $\frac{dQ}{dt}$

temperature gradient
potential gradient.

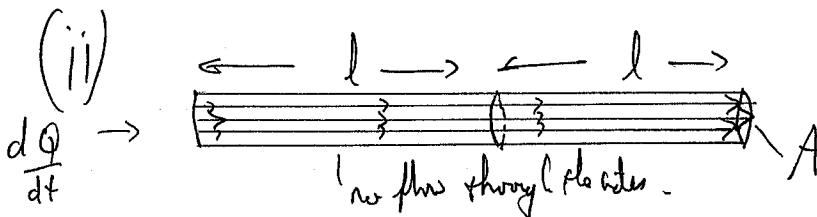
$$\frac{dQ}{dt} = -\lambda \cdot A \cdot \frac{\Delta T}{\Delta x}$$

$$I = \frac{V}{R} \Rightarrow \frac{dQ}{dt} = \frac{A}{l} \cdot (V_1 - V_2) \text{ or potential difference.}$$



$$\therefore \frac{dQ}{dt} = -\frac{l}{\rho} A \cdot \frac{dV}{dx}$$

ρ is the resistivity. $\frac{dV}{dx}$ is the potential gradient.



~~$\therefore A \cdot \Delta T = \Delta T_1 + \Delta T_2$ (temperatures add).~~

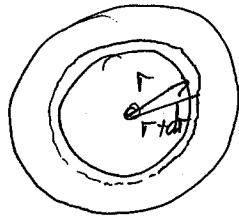
$$\therefore \frac{dQ}{dt} \frac{2l}{A} = \frac{dQ}{dt} \frac{l}{\lambda_1 A} + \frac{dQ}{dt} \frac{l}{\lambda_2 A}$$

(for the whole length.)

So that $\frac{2}{\lambda} = \frac{1}{\lambda_1} + \frac{1}{\lambda_2}$

Ques 1 cont.

(b)



Current flows through a thin "circumference" of thickness, \$dr\$, area \$2\pi r \times t\$ and resistivity \$\rho\$.

$$dR = \frac{\rho}{2\pi r} dt$$

$$\text{So } \int_0^R dR' = \int_{\pi r^2}^{\pi (r+dr)^2} \frac{dr'}{r}$$

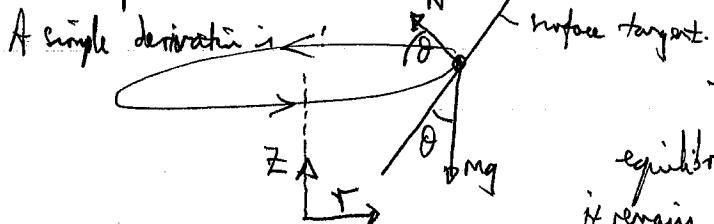
$$R = \frac{1}{2\pi r} \ln\left(\frac{r}{r_0}\right)$$

- (c) . The Wavefronts can be seen to be curved and bending slightly water - an effect of diffraction.
- There is a small acceleration of the water down the slope as the Wavefronts become a little further apart.
 - The Water is very shallow, so its speed is limited by friction with the ground. It can be seen to be shallow enough that the texture of the road shows in the ripples in the water surface (bottom left of Fig 1.2).
 - The limited increase of speed ensures that the wavefronts do not spread out much. But there is a noticeable pile-up at each Wavefront - as the water deepens slightly at a Wavefront, the surface water moves over the higher close to the ground, giving a rolling bridge effect appearance at the Wavefronts. This appears to maintain the Wavefront as it travels down the slope, so that they do not overtake each other and merge together.
 - There must be a slight dip in the road as the water moves towards the centre. This also limits the diffraction and may cause the curvature of the Wavefront, as the deeper water travels faster.

Ques 3

(a)

There are several ways to derive the shape of the surface by considering elements of the fluid.



The particle of mass m is not in equilibrium as it is accelerating inwards. But it remains at the same height up the slope.

$$\text{Resolving vertically: } N \sin \theta = mg$$

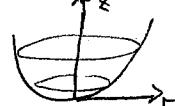
$$\text{horizontally: } N \cos \theta = m r \omega^2$$

$$\Rightarrow \tan \theta = \frac{mg}{mr\omega^2} = \frac{g}{r\omega^2}$$

$$\text{With } \theta \text{ specified, } \tan \theta = \frac{dr}{dz}$$

$$\text{So } \frac{dr}{dz} = \frac{g}{r\omega^2}$$

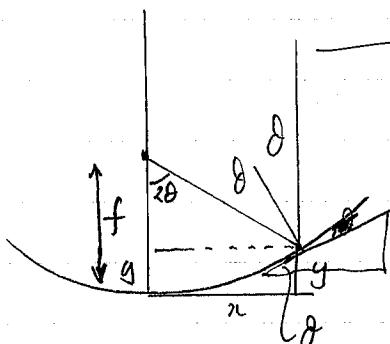
$$\text{Integrating, } \int r \omega^2 dr = \int g dz \Rightarrow z = \frac{r^2 \omega^2}{2g} + c$$



$$z=0 \text{ when } r=0 \Rightarrow c=0$$

$$\text{So } z = \frac{r^2 \omega^2}{2g} \text{ (parabola).}$$

(b)



$$\tan \theta = \frac{dy}{dx} = \frac{d(\alpha x^2)}{dx} = 2ax$$

Now trigonometry!

$$\sin \theta = 2 \sin \theta \cos \theta$$

$$\cos 2\theta = \cancel{(2 \cos^2 \theta - 1)}$$

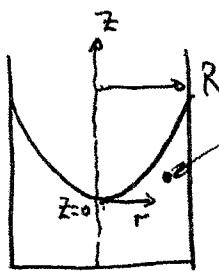
$$\therefore \tan 2\theta = \frac{2 \sin \theta \cos \theta}{2 \cos^2 \theta - 1}$$

$$= \frac{2 \tan \theta}{2 - \frac{1}{\cos^2 \theta}} = \frac{2 \tan \theta}{1 + (1 - \frac{1}{\cos^2 \theta})}$$

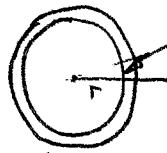
$$= \frac{2 \tan \theta}{1 - \tan^2 \theta} \quad \text{using } \tan \theta = 2ax$$

$$= \frac{4 \tan \theta}{1 - 4a^2 x^2}$$

Alternatives for parabolic surface of mercury



Consider point in the fluid which has an element of mass given by its small dimensions.

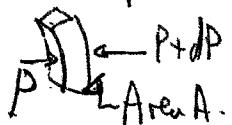


$$dm = r d\theta dr dz \rho$$

The element of mass is not in equilibrium since it is being accelerated.

The shape may be static but it is not an equilibrium shape.

From above:



$$A dP = dF \Rightarrow dF = dm \tau \omega^2$$

$$dP = \frac{dF}{A} = \frac{\tau dd dr dz \rho \tau \omega^2}{A \tau d\theta dz}$$

$$dP = \rho \omega^2 \tau dr$$

$$P = \rho \omega^2 \frac{r^2}{2} + c : r=0, P=0 \Rightarrow c=0$$

This answers that $v = \omega r$, i.e. the fluid flows is negligible. The fluid rotates over high unit.

(A)

If A a region in the fluid, B, $\rho \frac{r^2 \omega^2}{2} \rightarrow$

the pressures are:

$$\rho g(z - z_B) \quad \rho(r + dr)^2 \frac{\omega^2}{2}$$

$$\rho g(z - z_B + dz) = \rho(r + dr)^2 \frac{\omega^2}{2}$$

$$\text{and } \rho \frac{r^2 \omega^2}{2} = \rho g(z - z_B)$$

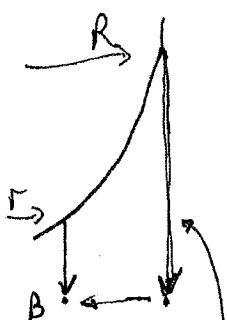
$$\therefore \rho g dz = \rho r dr \cdot \omega^2$$

$$\therefore gdz = \omega^2 r dr$$

$$z = \frac{\omega^2 r^2}{2g} + C$$

$$z=0, r=0 \Rightarrow C=0$$

(B)

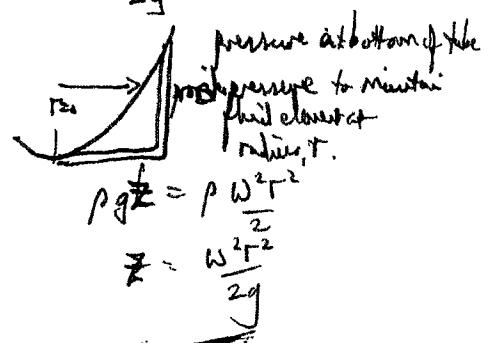


$$\rho g(z_R - z_B - (z - z_B)) = \rho \frac{R^2 \omega^2}{2} \rightarrow \tau \frac{r^2 \omega^2}{2}$$

$$(z_R - z) = \frac{\omega^2}{2g}(R^2 - r^2) \quad \text{at } z=0, r=R \Rightarrow z_R = \frac{\omega^2 R^2}{2g}$$

$$\therefore z = \frac{\omega^2 r^2}{2g}$$

(C)



Qn 3 cont.

(c) From the diagram we can see that $\tan 2\theta = \frac{2r}{f - y}$

To show that f is fixed, substitute for $y = a\sec^2\theta$

$$\text{Then } \tan 2\theta = \frac{2r}{f - a\sec^2\theta} = \frac{4ax}{4af - 4a^2x^2}$$

and so the term $4af = 1$

$$\text{if } f = \frac{1}{4a} \quad \text{which is constant for all paraxial rays.}$$

(d) From (a) we have $Z = \frac{\omega^2}{2g} r^2$ and $y = a\sec^2\theta$

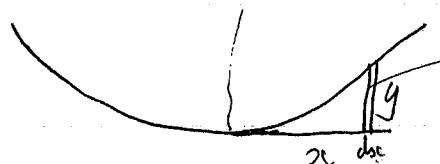
$$\text{so that } f = \frac{1}{4\omega^2/2g} = \frac{g}{2\omega^2}$$

$$D_{\min} = 1.22 \frac{\lambda}{D} \quad \text{using the Rayleigh criterion.}$$

The resolution has to be 40 times better for red light. (D will be better than this for blue).

$$\therefore D = 1.22 \frac{\lambda}{D_{\min}} = 1.22 \times \frac{700 \times 10^{-9}}{40 \times 10^{-3} \times \frac{1}{3600} \times \frac{\pi}{180}} \\ = 4.40 \text{ m}$$

$$\left[\text{and } f = \frac{9.81}{2 \times \left(\frac{8.5 \times 2\pi}{60} \right)^2} = \frac{6.19 \text{ m}}{\cancel{4.40 \text{ m}}} \right]$$



$$dV = 2\pi x \, dy \cdot y$$

$$= 2\pi x \, dy \cdot a \sec^2\theta$$

$$= 2\pi a x^3 \, dy$$

$$V = \int_0^{D/2} 2\pi a x^3 \, dy = 2\pi a \frac{x^4}{4} \Big|_0^{D/2}$$

$$= \frac{2\pi a D^4}{4 \cdot \frac{1}{4}} = \frac{\pi a D^4}{832}$$

$$\text{and as, } a = \frac{1}{4f}$$

$$V = \frac{\pi}{32f^4} D^4 \\ = \frac{0.1141938 \times 10^{-3}}{1.49 \text{ m}^3}$$

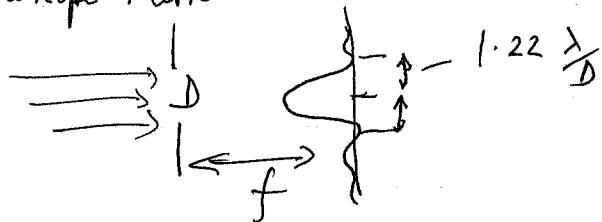
An 3 cont.

(e) Surface area $\sim \pi \frac{D^2}{4}$ = 15.2 m²

Rate of loss = 0.77 kg/week.

Rate is $\frac{0.77}{13600} \div 1.49 \times 100\% \text{ power.}$
 $= 3.8 \times 10^{-3}\% \text{ per week}$

(f) The image of the almost point source (the star) has a circular diffraction pattern in the focal plane of the telescope mirror.



$$r_{\text{diff}} = f \times 1.22 \frac{\lambda}{D}$$

$$\therefore \text{the intensity factor} = \frac{\pi(D/2)^2}{\pi r_{\text{diff}}^2} = \frac{D^2}{4f^2 1.22^2 \lambda^2}$$

$$= \frac{4 \cdot 40^2}{4 \times 6.19^2 \times 1.22^2 \times 700^2 \times 10^{-18}}$$
$$= 3.4 \times 10^{12}$$

$$\text{Intensity} = \frac{\text{Power}}{\text{Area}} = 3.4 \times 10^{12} \times 8.0 \times 10^{-14} \text{ Wm}^{-2}$$
$$= 0.27 \text{ Wm}^{-2}$$

averaged over the central maximum area.

Question 4. (a) $PV^\gamma = \text{const} = k$

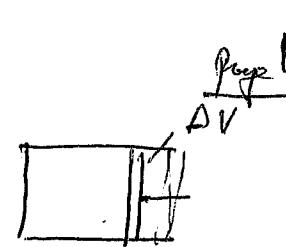
(Adiabatic) WD on a gas is $\frac{-PdV}{V^\gamma}$

$$= -k \frac{V_f^{\gamma-1}}{V_0^{\gamma-1}} \Big|_{V_0}^{V_f}$$

$$= -k \frac{(V_f^{1-\gamma} - V_0^{1-\gamma})}{1-\gamma}$$

$$= \cancel{k(V_f^{1-\gamma} - V_0^{1-\gamma})}$$

$$= k \frac{(V_f^{1-\gamma} - V_0^{1-\gamma})}{\gamma-1}$$

$$k = P_0 V_0^\gamma = P_f V_f^\gamma$$


WD of gas is $PΔV$

(Adiabatic) WP on a gas $= \frac{(P_f V_f - P_0 V_0)}{(\gamma-1)}$.

WD on the gas on the RHS. is given by

$$WD = \frac{(P_f V_f - P_0 V_0)}{(\gamma-1)} \quad \text{with} \quad P_f = \frac{27}{8} P_0$$

$$\gamma = 1.5$$

and V_f given by

$$P_0 V_0^\gamma = P_f V_f^\gamma$$

$$\text{so that, } P_0/V_0^\gamma = \frac{27}{8} P_0 \cdot V_f^\gamma$$

$$\text{Hence, } \left(\frac{V_0}{V_f}\right)^{1.5} = \left(\frac{27}{8}\right)$$

$$\begin{aligned} \frac{V_0}{V_f} &= \frac{9}{4} \\ V_f &= \frac{4}{9} V_0 \end{aligned}$$

$$\left. \text{and } P_0 V_0 = n R T_0 \right\}$$

This result can also be obtained by knowing T_f from part (b) ($T_f = \frac{3}{2} T_0$). However care must be taken as the molecules are not point particles so $U_0 = 2NkT_0$, not $\frac{3}{2}kT_0$. With $\gamma = \frac{3}{2}$, $\gamma = (1 + \frac{2}{f})$ with the degrees of freedom. Hence, here, $f = 4$ (3 translation and 1 other).

$$\text{so } U_0 = 4 \cdot \frac{1}{2} N k T_0 \quad \text{and with } T_f = \frac{3}{2} T_0$$

$$U_f = 4 \cdot \frac{1}{2} N k \left(\frac{3}{2} T_0\right) \quad \text{so } \Delta U = 2 N k \cdot \frac{3}{2} T_0 - 2 N k T_0 = \underline{\underline{n R T_0}}$$

Question 8 cont. ---

$$(b) PV^\gamma = \text{const. } ①$$

$$\text{and } \frac{PV}{T} = \text{const} \text{ so that } \frac{PV^\gamma}{T^\gamma} = \text{const. } ②$$

Hence, dividing ① by ②

We obtain $P^\frac{1}{\gamma-1} \cdot T = \text{const}$

Now we can substitute.

$$P_0^{\frac{2}{\gamma-1}} \cdot T_0 = \left(\frac{27}{8} P_0\right)^{\frac{2}{\gamma-1}} \cdot T_f$$

$$\frac{T_0}{P_0^{\frac{1}{\gamma-1}}} = \left(\frac{27}{8} P_0\right)^{-\frac{1}{3}} \cdot T_f$$

$$\frac{T_0}{P_0^{\frac{1}{\gamma-1}}} = \frac{2}{3} T_f$$

$$T_0 \approx \frac{2}{3} T_f$$

$$\text{so } T_f = \underline{\underline{\frac{3}{2} T_0}}$$

Alternatively

$$dU = nC_V \cdot \Delta T$$

and we know from part (a) that

$dU = \text{the WD on the gas}$

even though the volume is not constant
here. The internal energy change is due
to the temperature rise and the molar specific
heat capacity when no work is done or
by the gas.

in an adiabatic compression
in which no heat is exchanged with the surroundings.

$$\therefore nC_V(T_f - T_0) \leq nRT$$

$$T_f - T_0 = \frac{R}{C_V} \cdot T_0$$

$$T_f - T_0 = (\gamma - 1)T_0$$

$$T_f = \gamma T_0 - T_0 + T_0$$

$$T_f = \underline{\underline{\frac{3}{2} T_0}} \text{ or } T_f = \underline{\underline{\left(1 + \frac{R}{C_V}\right) T_0}}$$

Now this step needs

Mayer's equation

$$C_p - C_V = R$$

$$\text{so that } \frac{C_p}{C_V} - 1 = \frac{R}{C_V}$$

$$\gamma - 1 = \frac{R}{C_V}$$

(c) Pressure is $\frac{27}{8} P_0$ on LHS, same as on RHS.

And the total volume of $2V_0$ is $V_{LHS} + V_{RHS}$
and we know from part(a) that $V_{RHS} = \frac{4}{9} V_0$

$$\text{Hence } V_{LHS} = 2V_0 - \frac{4}{9} V_0$$

$$V_{LHS} = \frac{14}{9} V_0$$

~~or alternatively~~ Details: $P_0 V_0^\gamma = P_f V_f^\gamma$ for the RHS.

$$P_0 V_0^\gamma = \frac{27}{8} P_0 \times V_f^\gamma$$

$$\text{RHS} \rightarrow V_f = \left(\frac{8}{27}\right)^{\frac{1}{\gamma}} V_0 = \frac{4}{9} V_0$$

$$\text{Hence. } \text{LHS} \rightarrow V_f = \frac{14}{9} V_0$$

Using the ideal gas law.

$$\text{for LHS, } \frac{P_0 V_0}{T_0} = \frac{P_f V_f}{T_f}$$

$$\frac{P_0 V_0}{T_0} = \frac{\frac{27}{8} P_0 \times \frac{14}{9} V_0}{T_f}$$

$$T_f = \frac{27}{8} \cdot \frac{14}{9} T_0$$

$$T_f = \underline{\underline{\frac{21}{4} T_0}}$$

- (d) The gas on the LHS has thermal energy supplied, so it gains internal energy as its temperature rises, but it also does work in expanding. Remember, as thermal energy is supplied, this is not an adiabatic change, so the adiabatic formula for WD does not apply on the LHS.

$$\begin{aligned}\text{Gain in internal energy, } \Delta U &= n C_V (T_f - T_0) \\ &= n C_V \left(\frac{21}{4} T_0 - T_0 \right) \\ &= n \cdot C_V \cdot \frac{17}{4} T_0\end{aligned}$$

For if we are $C_P - C_V = R$ so that $\gamma - 1 = \frac{R}{C_V}$ then

$$\begin{aligned}\Delta U &= n R \frac{17}{4} T_0 \\ &= \underline{\underline{\frac{17}{2} P_0 V_0}}\end{aligned}$$

Since the WD on the RHS is $P_0 V_0$ or $n R T_0$ from part (a)
this is the work done by the gas on the LHS.

$\therefore \Delta Q = \text{increase in thermal energy} + \text{work done by the gas in expanding}$

$$\begin{aligned}&= \underline{\underline{\frac{17}{2} P_0 V_0}} + P_0 V_0 \\ &= \underline{\underline{\frac{19}{2} P_0 V_0}} = \underline{\underline{\frac{19}{2} n R T_0}} = n C_V \frac{17}{4} T_0 + n R T_0 \\ &\quad = \underline{\underline{n T_0 \left(\frac{17}{4} C_V + R \right)}}.\end{aligned}$$

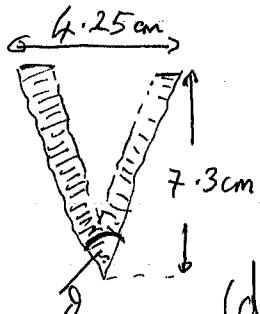
Question 5

(a) Scale : on paper 10.2 cm corresponds to 7.8 cm on screen.

$$1 \text{ cm} \quad " \quad \therefore 0.765 \text{ cm on screen}$$

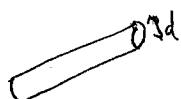
(i) Smallest fringe spacing : 24 fringes in 5.2 cm is 0.166 cm on screen

(ii) Highest spacing 4 fringes (diagonal) in 4.2 cm is 0.803 cm on screen
 (and there are 5 small fringes in 1 large fringe)



$$\theta = 32.5^\circ = 33^\circ$$

(d) The wire thickness produces the wide spaced fringe pattern.



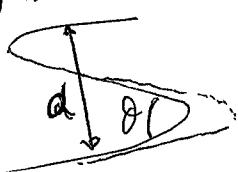
$$\frac{\lambda}{\text{Wire diameter}} = \frac{\text{fringe spacing}}{\text{distance to screen}}$$

$$\frac{\lambda}{d} = \frac{W}{D}$$

$$\therefore d = \frac{\lambda D}{W} = \frac{633 \times 10^{-9} \times 4.2}{0.803 \times 10^{-2}} \\ = 3.3 \times 10^{-4} \text{ m} \\ = 0.33 \text{ mm.}$$

N.B. The diffraction pattern does not correspond to the dimension of the grating illustrated.

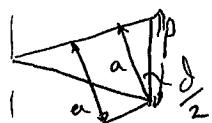
(iii)



$$a = \frac{\lambda D}{W} = \frac{633 \times 10^{-9} \times 4.2}{0.166 \times 10^{-2}} \\ = 1.60 \times 10^{-3} \text{ m}$$

$$= 1.6 \text{ mm} \\ \underline{= 1.67} \\ = 1.7 \text{ mm}$$

(iv)



Pitch, $p = \frac{a}{\cos \alpha}$

$$\frac{2\pi r}{p} = \tan(90^\circ - \alpha)$$

(v)



projecting is a sine wave

$$2r, 2r = 1.79 \text{ mm}, r = 0.897 \text{ mm}$$

Ques. 5 cont.

(b) (i) Scale 9.1 cm on paper is 9.4 cm on screen
4.5 cm for 10 fringes
 \therefore fringe on screen is $W = 0.465 \text{ cm}$

$$(ii) a = \frac{\lambda D}{W} = \frac{0.15 \times 10^{-9} \times 9.0 \times 10^{-2}}{0.465 \times 10^{-2}} \\ = 2.9 \times 10^{-9} \text{ m} \\ = \underline{2.9 \text{ nm}}$$

$$(iii) \theta = 84^\circ$$

$$(iv) \text{pitch, } p = \frac{a}{\sin \frac{\theta}{2}} = \frac{2.9}{\sin 42^\circ} = \underline{3.9 \text{ nm}}$$

$$(v) \text{radius; } \frac{2\pi r}{\lambda} = \tan(90 - \frac{\theta}{2})$$

$$r = \underline{0.69 \text{ nm}}$$