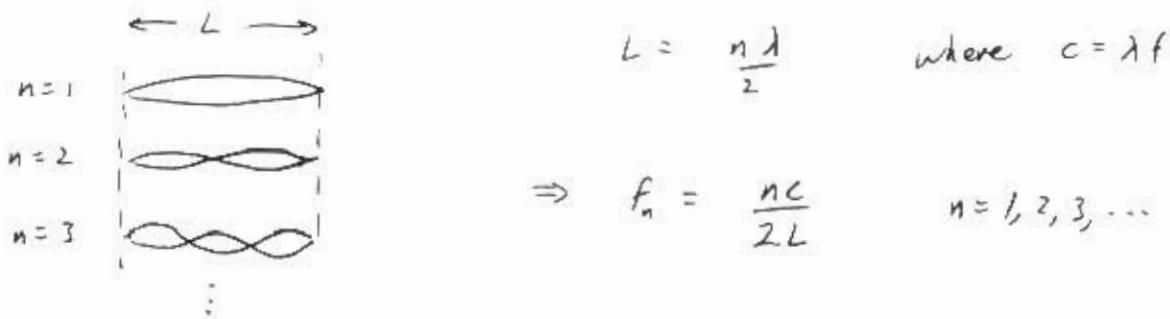


(a) The modes are



$$L = \frac{n\lambda}{2} \quad \text{where } c = \lambda f$$

$$\Rightarrow f_n = \frac{nc}{2L} \quad n = 1, 2, 3, \dots$$

$$(b) n_0 = \frac{2Lf_{n_0}}{c} = \frac{2(1.5\text{ m})(5 \times 10^{14} \text{ /s})}{3 \times 10^8 \text{ m/s}} = 5 \times 10^6$$

$$(c) f_{n_0} \pm \Delta f = (n_0 \pm \Delta n) \frac{c}{2L}$$

$$f_{n_0} \pm \Delta f = \frac{n_0}{2L} \pm \frac{c}{2L} \Delta n$$

$$\Delta n = \frac{2L}{c} \Delta f = \frac{2(1.5\text{ m})(1 \times 10^9 \text{ /s})}{3 \times 10^8 \text{ m/s}} = 10$$

so the frequencies being excited into resonance are those corresponding to $n_0, n_0 \pm 1, n_0 \pm 2, \dots, n_0 \pm 10$

21 modes

(d) From (c), we note that as L gets shorter, Δn gets smaller (Δf is fixed by the plasma tube, not the cavity). Decreasing L , the first modes to go away are $n_0 \pm 10$, next is $n_0 \pm 9$, etc. To top off ± 10 modes, the length L must be $\frac{1}{10}$ as long as before (since $L \propto \Delta n$). Hence

$$L = \frac{1.5\text{ m}}{10} = 15\text{ cm}$$

Each hop is a projectile motion, so that

$$x = v_{ox} t, \quad (2)$$

and

$$t = 2 v_{oy} / g \quad (\text{from } \theta = y = v_{oy} t - \frac{1}{2} g t^2), \quad (2)$$

and upon eliminating t , the range L of a hop is

$$L = 2 v_{ox} v_{oy} / g. \quad (2)$$

Since $v_{ox,n+1} = \epsilon_y^n v_{oy,n}$ and $v_{oy,n+1} = \epsilon_x^n v_{oy,n}$, it follows that

$$v_{ox,n+1} = \epsilon_y^n v_{oy,n} \quad (1)$$

$$v_{oy,n+1} = \epsilon_x^n v_{oy,n} \quad (1)$$

$$L_{n+1} = (\epsilon_x \epsilon_y)^n L_n \quad (2)$$

$$t_{n+1} = L_{n+1} / v_{oy,n+1} = \epsilon_y^n t_n \quad (2)$$

The total range is

$$\begin{aligned} L^* &= L_1 + L_2 + L_3 + L_4 + \dots \\ &= L_1 [1 + \epsilon_y \epsilon_x + (\epsilon_y \epsilon_x)^2 + (\epsilon_y \epsilon_x)^3 + \dots] \\ &= L_1 [1 + \epsilon_y \epsilon_x]^{-1} \end{aligned} \quad (2)$$

The total time is

$$\begin{aligned} t^* &= t_1 + t_2 + t_3 + t_4 + \dots \\ &= t_1 [1 + \epsilon_y + (\epsilon_y)^2 + (\epsilon_y)^3 + \dots] \\ &= t_1 [1 - \epsilon_y]^{-1} \end{aligned} \quad (2)$$

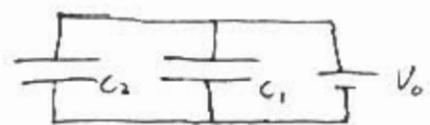
The angle θ_i is given by

$$\begin{aligned} \theta_i &= \tan^{-1} (v_{oy,i} / v_{ox,i}) \\ &= \tan^{-1} (g t_i^2 / 2 L_i) \\ &= \tan^{-1} \{[g t^{*2} (1 - \epsilon_y)^2] / [2 L^* (1 - \epsilon_y \epsilon_x)]\} \end{aligned} \quad (4)$$

(a) as $t \rightarrow \infty$ The circuit becomes:

In parallel, so

$$V_1 = V_2 = V_o = 10V$$

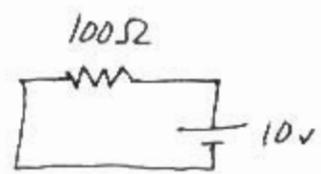


$$\Rightarrow q_1 = C_1 V_o = (1\mu F)(10V) = 10\mu C$$

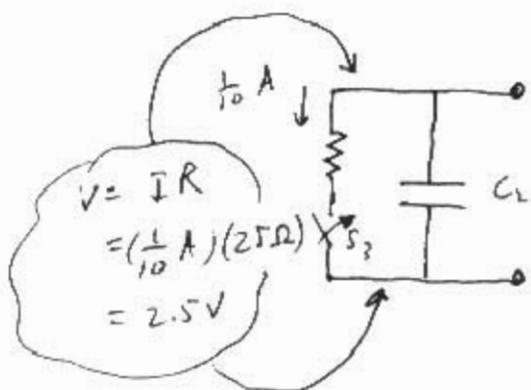
$$q_2 = C_2 V_o = (4\mu F)(10V) = 40\mu C$$

(b) as $t \rightarrow \infty$, current flows through outer loop only,

$$I = \frac{10V}{100\Omega} = \frac{1}{10} A$$

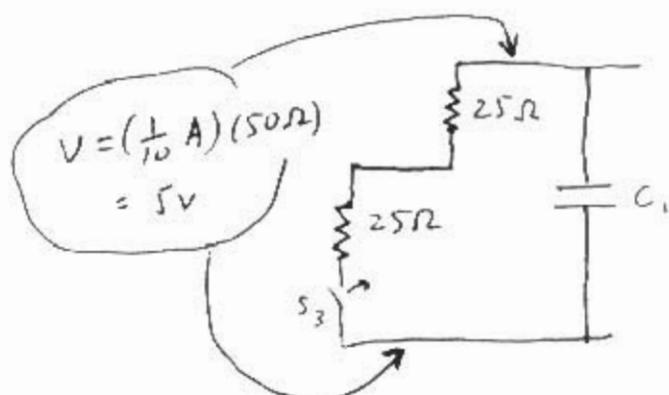


C_2 is in parallel with the 25Ω resistor by S_3 :



$$\Rightarrow q_2 = C_2 V_2 = (4\mu F)(2.5V) = 10\mu C$$

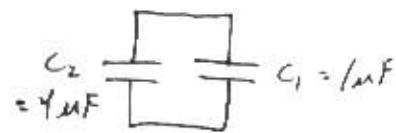
C_1 is in parallel with 50Ω



$$q_1 = C_1 V_1 = (1\mu F)(5V) = 5\mu C$$

$$(c) \text{ From (b), } q_{\text{total}} = q_1 + q_2 = 15 \mu\text{C}$$

Now the charge may only redistribute itself until $V_1 = V_2$



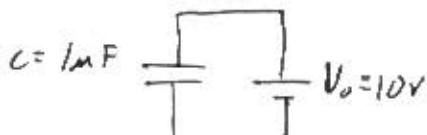
$$q_1 + q_2 = 15 \mu\text{C} \quad \text{and} \quad \frac{q_1}{C_1} = \frac{q_2}{C_2}$$



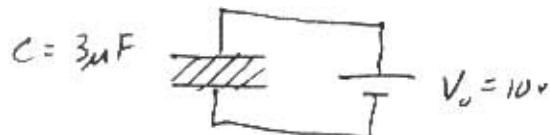
$$\leftarrow \quad q_1 = \frac{C_1}{C_2} q_2 = \frac{1}{4} q_2$$

$$\Rightarrow \frac{1}{4} q_2 + q_2 = 15 \mu\text{C} \quad \text{or} \quad q_2 = 12 \mu\text{C} \quad \Rightarrow \quad q_1 = 3 \mu\text{C}$$

(d)

initial

$$q_i = (1 \mu\text{F})(10\text{V}) = 10 \mu\text{C}$$

final

$$q_f = (3 \mu\text{F})(10\text{V}) = 30 \mu\text{C}$$

Battery connected, V does not change, but q does

$$\text{Work} = \Delta(\text{Energy})$$



$$W_{\text{Battery}} + W_{\text{sys}} = \Delta(\frac{1}{2}CV^2)$$



$$= \frac{1}{2}(3C_1 - C_1)V^2$$

$$= \frac{1}{2}(2 \mu\text{F})(10\text{V})^2$$

$$= 100 \mu\text{J}$$

(a) To conserve momentum the atom must recoil.

Thus after emitting the photon, the atom carries some kinetic energy K , leaving less than E_0 for photon.

(b)

Before decay

 (M')

$$(\text{Energy, momentum}) = (M'c^2, 0)$$

After decay

 $\leftarrow (M)$

$$(Mc^2 + K, -p)$$

 $\xrightarrow{\text{photon}}$

$$(E, E/c)$$

Conservation of Energy

$$M'c^2 = Mc^2 + K + E$$

$$(M' - M)c^2 = K + E$$

$$E_0 = K + E$$

Given $\frac{E_0}{2mc^2} \ll 1$

since K and E are $\ll E_0$,

$$\frac{K}{2mc^2} \text{ and } \frac{E}{2mc^2} \text{ are } \ll 1$$

too

$$0 = \frac{E}{c} - p$$

$$p = \frac{E}{c}$$

Since the atom is heavy,
 $v \ll c$ and $K = p^2/2m$, or
 $p \approx \sqrt{2mK}$ for atom

$$\sqrt{2mK} = \frac{E}{c} \Rightarrow K = \frac{E^2}{2mc^2}$$

{ or, relativistically,

$$(Mc^2)^2 = (K + Mc^2)^2 - (pc)^2$$

$$\Rightarrow K = Mc^2 \left[\sqrt{1 + \left(\frac{E}{Mc^2} \right)^2} - 1 \right] \quad \{$$

$$\approx \frac{E^2}{2mc^2}$$

Conservation of Energy Equation becomes

$$E_0 = \frac{E^2}{2mc^2} + E$$

$$\Rightarrow E = E_0 \left(1 + \frac{E}{2mc^2} \right)^{-1}$$

$$\approx E_0 \left(1 - \frac{E}{2mc^2} \right) \quad \text{Iterate}$$

$$= E_0 \left[1 - \frac{E_0 \left(1 - E/2mc^2 \right)}{2mc^2} \right]$$

Alternate: Quadratic formula -

$$E^2 + YE + \delta E_0 = 0$$

$$\text{where } \delta = 2mc^2$$

$$\Rightarrow E = \frac{\delta}{2} \left[-1 \pm \left(1 + \frac{4E_0}{\delta} \right)^{\frac{1}{2}} \right]$$

$$= \frac{\delta}{2} \left[-1 \pm \left(1 + 2\frac{E_0}{\delta} - \frac{E_0^2}{\delta^2} \right)^{\frac{1}{2}} \right]$$

since $E > 0$ take +

$$E = E_0(1 - \gamma)$$

Note: 2nd order needed
in Binomial

$$\text{If } \gamma \equiv \frac{E_0}{2mc^2} \ll 1 \quad \text{Then} \quad \frac{EE_0}{(2mc^2)^2} \approx \gamma^2 \text{ is negligible}$$

$$\Rightarrow E \approx E_0 \left(1 - \frac{E_0}{2mc^2} \right)$$

(c) From part (b),

$$E_0 - E = \frac{E_0^2}{2mc^2}$$

Using $E = hf$, and writing $E_0^2 = h f_0 E_0$,

$$h(f_0 - f) = \frac{hf_0 E_0}{2mc^2}$$

$$\frac{\Delta f}{f_0} = \frac{E_0}{2mc^2} = \frac{4 \times 10^5 \text{ eV}}{2(2 \times 10^6 \text{ eV})} = 10^{-6}$$

(d) For $v \ll c$, Doppler shift says

$$\frac{v}{c} \approx \frac{\Delta\lambda}{\lambda} \quad \text{where } c = \lambda f$$

Since $f_0 > f$, The observer must move towards the photon



$$\Delta\lambda = c \Delta\left(\frac{1}{f}\right) = -\frac{\Delta f}{f^2} c$$

\therefore (in magnitude)

$$\frac{v}{c} \approx \frac{\Delta f}{f^2} \frac{c}{\lambda}$$

$$= \frac{\Delta f}{f}$$

$$v = \frac{\Delta f}{f_0} c$$

$$= 10^{-6} c \quad (\text{using } \frac{\Delta f}{f_0} = 10^{-6} \text{ from part (c)})$$

$$= 300 \text{ m/s}$$

Alternate:

$$c = \lambda f$$

$$dc = d\lambda f + \lambda df$$

$$d\lambda = -\lambda \frac{df}{f}$$

$$\frac{d\lambda}{\lambda} = -\frac{df}{f}$$

minus sign means that a decrease in λ corresponds to increase in frequency!

(a) Use Gauss' law:

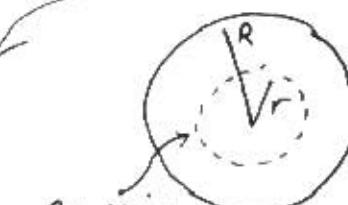
$$\text{flux of } \vec{E} = \frac{1}{\epsilon_0} (\text{charge enclosed})$$

↙

$$E_{\text{z axis}} = \frac{1}{\epsilon_0} (n_0 e) (\pi r^2 h)$$

$$E = \left(\frac{n_0 e}{2\epsilon_0} \right) r$$

$$\vec{E} = \frac{n_0 e}{2\epsilon_0} r (-\hat{r}) \quad \begin{matrix} \text{radially} \\ \text{in} \end{matrix}$$



Gaussian surface: cylinder,
 $r \leq R$
height h

no flux through end caps
($\vec{E} \cdot \hat{n} = 0$ There)

$$(b) \vec{F} = q(\vec{E} + \vec{v} \times \vec{B})$$

$$= -e \left[-\frac{n_0 e r \hat{r}}{2\epsilon_0} + rwB_0 \hat{r} \right]$$

$$\vec{B} = B_0 \hat{k}$$

$$\vec{v} = rw\hat{\theta}$$

$$\hat{\theta} \times \hat{k} = \hat{r}$$

$$\vec{F} = \left(\frac{n_0 e^2 r}{2\epsilon_0} - rwB_0 e \right) \hat{r}$$

$$(c) \vec{F} = m\vec{a}$$

$$\left(\frac{n_0 e^2 r}{2\epsilon_0} - rwB_0 e \right) \hat{r} = m \left(\frac{rw}{r} \right)^2 (-\hat{r})$$

$$\Rightarrow \frac{n_0 e^2}{2\epsilon_0} - wB_0 e + mw^2 = 0$$

$$w^2 - \left(\frac{B_0 e}{m} \right) w + \frac{n_0 e^2}{2\epsilon_0 m} = 0$$

$$w^2 - w_c w + \frac{1}{2} w_p^2 = 0$$

$$w_c = \frac{B_0 e}{m}$$

$$w_p^2 = \frac{n_0 e^2}{\epsilon_0 m}$$

$$\omega = \frac{\omega_c}{2} \pm \frac{1}{2} \sqrt{\omega_c^2 - 2\omega_p^2}$$

$$\omega = \frac{\omega_c}{2} \left(1 \pm \sqrt{1 - 2\left(\frac{\omega_p}{\omega_c}\right)^2} \right)$$

(d) Since ω is real, this requires

$$1 - 2\left(\frac{\omega_p}{\omega_c}\right)^2 \geq 0$$

$$\omega_p^2 \leq \frac{1}{2} \omega_c^2$$

$$\frac{n_0 e^2}{\epsilon_0 m} \leq \frac{1}{2} \frac{e^2 B_0^2}{m^2}$$

$$n_0 \leq \frac{B_0^2 \epsilon_0}{2m} \quad \text{so} \quad (n_0)_{\max} = \frac{B_0^2 \epsilon_0}{2m}$$

$$(e) \quad (n_0)_{\max} = \frac{B_0^2 \epsilon_0}{2m} \quad \text{use} \quad \mu_0 \epsilon_0 = \frac{1}{c^2}$$

$$= \frac{B_0^2}{2m(\mu_0 c^2)}$$

$$= \frac{B_0^2 / 2\mu_0}{mc^2} = \frac{\text{magnetic field energy density}}{\text{electron mass energy}}$$

$$\oint \vec{B} \cdot d\vec{l} = \mu_0 I_s$$

where S = boundary of S

$$B' h = \mu_0 \int j da$$

where j = current density

$$= (n_0 e) v$$

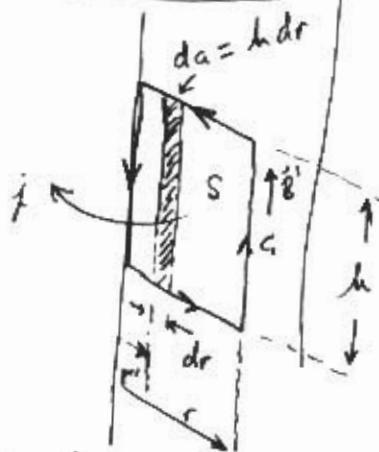
$$= n_0 e r' w$$

$$\text{or } \sum r' dr$$

$$B' h = \mu_0 (n_0 e w) \int_0^r r' (dr' h)$$

$$B' h = \mu_0 n_0 e w h \int_0^r r' dr'$$

$$B' = \mu_0 n_0 e w \frac{r^2}{2}$$



$$\vec{B}' = \text{induced field} = B' \hat{k}$$

given $\vec{B}' = 0$ on axis

Note $\vec{B}' \cdot d\vec{l} = 0$ horizontally

(f) Force due to induced magnetic field is \vec{F}' where

$$\vec{F}' = -e \vec{v} \times \vec{B}' = -e(rw\hat{\theta}) \times (\mu_0 n_0 e w \frac{r^2}{2} \hat{k})$$

$$= \frac{\mu_0 n_0 e^2 w^2 r^3}{2} (-\hat{r})$$

$$v = rw$$

$$\hat{\theta} \times \hat{k} = \hat{r}$$

$$|\vec{F}'| = \frac{1}{2} \mu_0 n_0 e^2 v^2 r$$

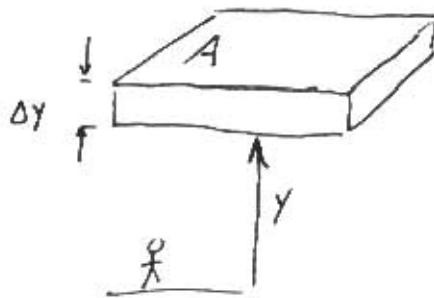
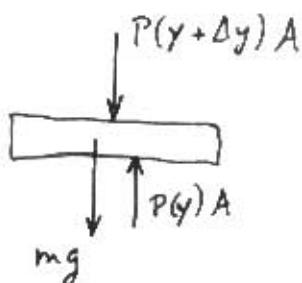
$$\therefore \frac{|\vec{F}'|}{|\vec{F}_{el.}|} = \frac{\frac{1}{2} \mu_0 n_0 e^2 v^2 r}{\frac{1}{2} \frac{n_0 e^2 r}{\epsilon_0}} = \mu_0 \epsilon_0 v^2 = \frac{v^2}{c^2}$$

$|e\vec{E}|$ where $|\vec{E}|$ is from part (a)

- (a) Apply $F = ma$ to a slab of atmosphere of thickness Δy , cross-sectional area A , density ρ . Take y positive upwards.

geometry:

Forces:



$$F = ma \quad (\text{which} = 0)$$

$$P(y)A - P(y+\Delta y)A - (\rho A \Delta y)g = 0$$

$$\underbrace{P(y+\Delta y) - P(y)}_{\Delta P} = -\rho g \Delta y$$

$$\Delta P = -\rho g \Delta y$$

note $\rho = \rho(y)$
 $g = g(y)$

(b) Since $v_o = -\frac{\Delta y}{\Delta t}$ (probe going down, y points up) $\frac{v_o}{g}$

$$\Delta P = -\rho g \Delta y \quad \text{gives}$$

$$\frac{\Delta P}{\Delta t} = \rho g v_o \quad \dots (1)$$

Need ρ : From ideal gas eq. of state $PV = nRT$,

$$P = \frac{n}{V} RT \quad \text{where } \frac{n}{V} = \frac{\# \text{ moles}}{\text{vol.}}$$

↗

$$= \frac{\# \text{ moles}}{\text{mass}} \frac{\text{mass}}{\text{vol.}}$$

$$\frac{n}{V} = \frac{1}{M} \rho$$

↙

$$\Rightarrow P = \frac{\rho}{M} RT$$

$$\rho = \frac{PM}{RT} \quad \dots (2)$$

using ρ from (2) in (1),

$$\frac{\Delta P}{\Delta t} = \frac{PM}{RT} g v_o$$

$$v_o = \left(\frac{\Delta P}{\Delta t} \right) \frac{RT}{PMg} \quad \dots (3)$$

B/C
3/1

From graph, probe lands on surface at $t = 4000 \text{ s}$

$$P = 60 \text{ units}$$

$$g = 9.9 \text{ N/kg}$$

$$T = 400 \text{ K}$$

From graph: at surface

$$\frac{\Delta P}{\Delta t} = \frac{(60 - 55) \text{ units}}{(4000 - 3900) \text{ s}} = .05 \text{ units/s}$$

From Eq. (3),

[line from $(3900, 55)$ to $(4000, 60)$ is straight]

$$v_0 = \frac{(.05 \text{ units/s})(8.3 \text{ J/K-mol.})(400 \text{ K})}{(60 \text{ units})(44 \times 10^{-3} \text{ kg/mol})(9.9 \text{ N/kg})} \approx 6.4 \text{ m/s}$$

(c) Probe was 15 km above surface

$$\Delta t = \frac{15 \text{ km}}{6.4 \text{ m/s}} \approx 2344 \text{ s} \text{ before touchdown; i.e., at}$$

$$t = (4000 - 2344) \text{ s} = 1656 \text{ s}$$

at $t \approx 1650 \text{ s}$, $P = 11 \text{ units}$

slope at $t = 1650 \text{ s}$ is about the same as slope
of line from $(1500 \text{ s}, 9 \text{ units})$ to $(2000 \text{ s}, 15 \text{ units})$

$$\text{slope} \approx \frac{(15 - 9) \text{ units}}{(2000 - 1500) \text{ s}} = .012 \text{ units/s}$$

$\approx .01$ to two sig. figures (like slope
found above)

From Eq. (3),

$$T = \frac{PMg v_0}{R \left(\frac{\Delta P}{\Delta t} \right)}$$

B2

4/4

$$\Rightarrow \frac{T_2}{T_1} = \frac{P_2}{P_1} \frac{(\Delta P/\Delta t)_1}{(\Delta P/\Delta t)_2}$$

point 1: surface
Point 2: 15 km above surface

$$T_2 = \frac{11}{60} \left(\frac{.05}{.01} \right) 400 \text{ K} \approx 367 \text{ K}$$

(d) $g(h) = \frac{GM}{(R+h)^2}$

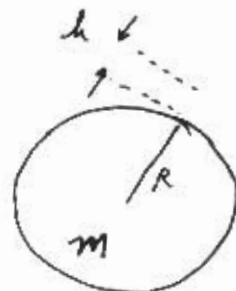
$$= \frac{GM}{R^2} \left(1 + \frac{h}{R} \right)^{-2}$$

$$\approx g_0 \left(1 - \frac{2h}{R} \right)$$

$$\equiv g_0 - \Delta g$$

where $\Delta g = \frac{2h}{R} g_0$

$$\frac{\Delta g}{g_0} = \frac{2h}{R} = \frac{2(15 \times 10^3 \text{ m})}{5 \times 10^6 \text{ m}} = 6 \times 10^{-3}$$



$$g_0 = \frac{GM}{R^2}$$

In part (c), The estimate of P and $\Delta P/\Delta t$ at 15 km above the surface was much cruder than 6 parts out of a thousand.